Overview of Particle Physics

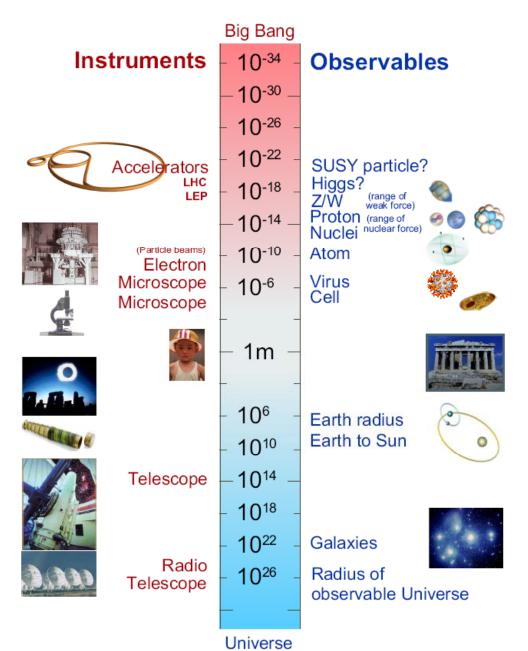
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Quarknet 2008 at UCR

Particle Physics

- What is it?
 - Study of the elementary constituents of matter
 - And the fundamental forces that control their behavior at the most basic level
 - Connections to cosmology, i.e. study of the universe on the largest scales

The Size of Things

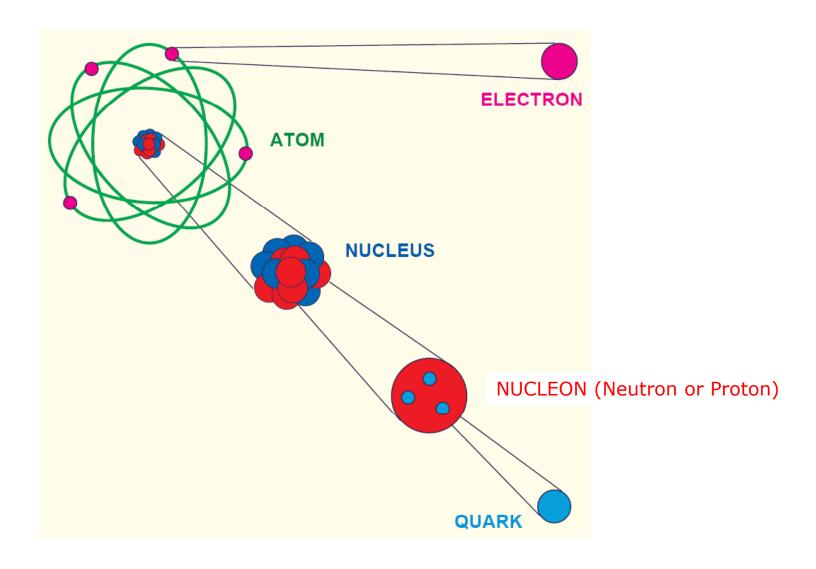


- Wavelength of probe radiation should be less than size of object probed
- de Broglie wavelength

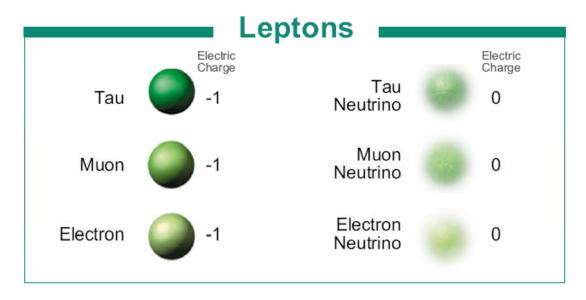
$$\lambda = \frac{h}{p} = \frac{hc}{E}$$

i.e. high energy probes resolve smaller sizes

Structure of the Atom



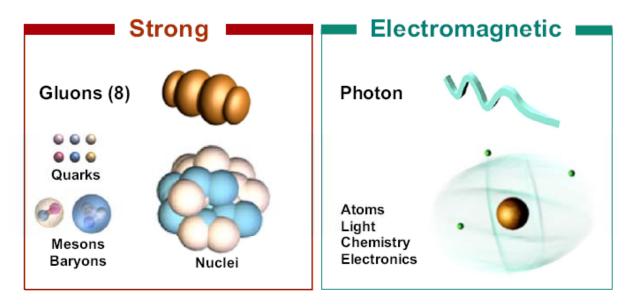
Elementary Particles

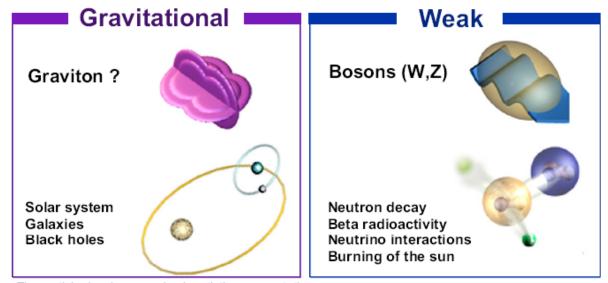




The particle drawings are simple artistic representations

Forces



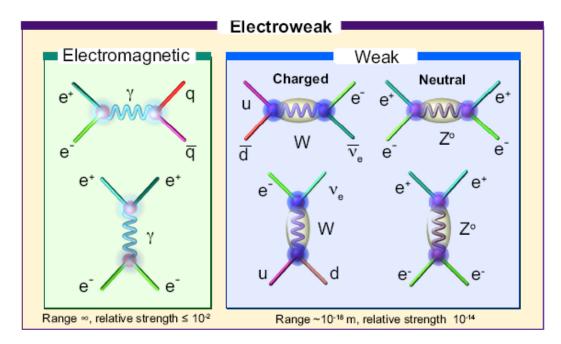


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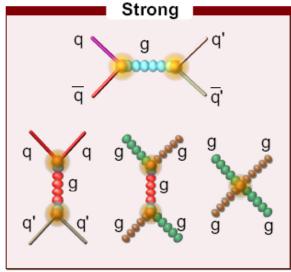
Interactions: Coupling of Forces to Matter

- Newton's law:
 - Action at a distance

- Maxwell's equations:
 - Interaction by fields



- Quantum Field Theory:
 - Forces are produced by the exchange of "messengers" or "gauge bosons"

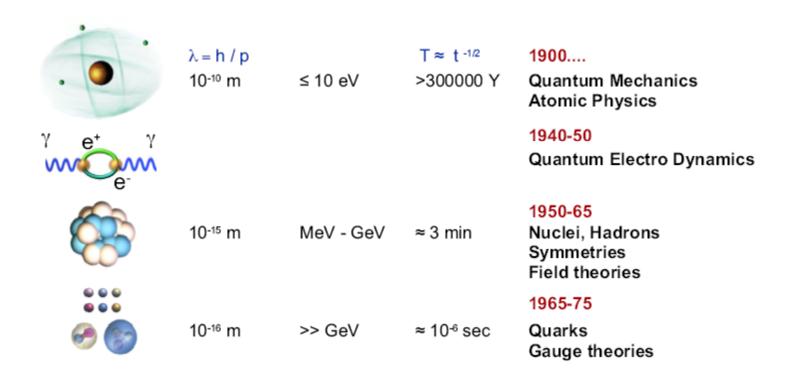


Range ~ 10-15 m, relative strength = 1

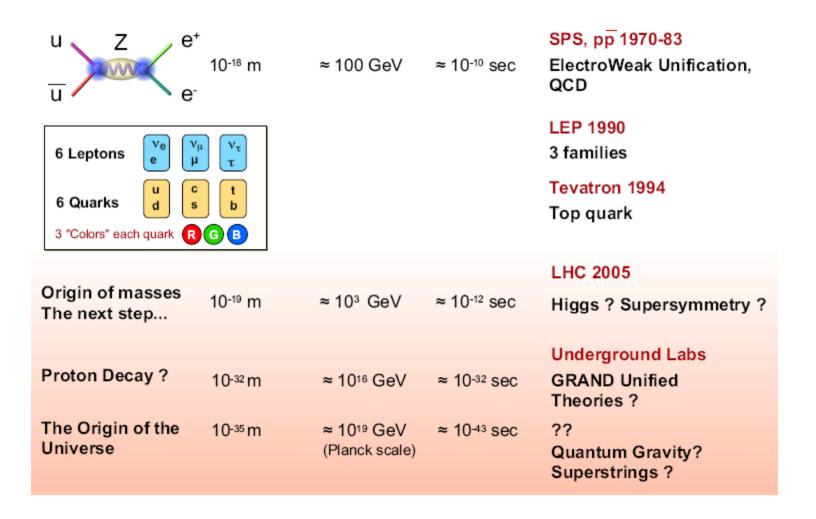
Aside: Units of Energy and Mass

- Physicists usually express energy in units of electron volts (eV), where one eV is the energy accquired by an electron accelerated through a potential of 1 Volt
- $1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$
- Masses are usually expressed in terms of eV/c^2 using Einstein's relation $F = m c^2$
 - Example: Rest mass of electron
 - Use $E = m c^2 = (9.11 \times 10^{-32} \text{ kg}) (3 \times 10^8 \text{ m/s}^2) = 8.2 \times 10^{-14} \text{ J}$
 - Convert to eV: E = 0.511 MeV
 - So, $m_e = 0.511 \text{ MeV}/c^2$

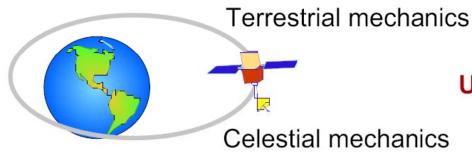
More on Sizes



More on Sizes

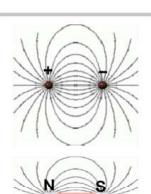


Unification of the Forces



Universal Gravitation

Inertial vs. Gravitational mass (I. Newton, 1687)



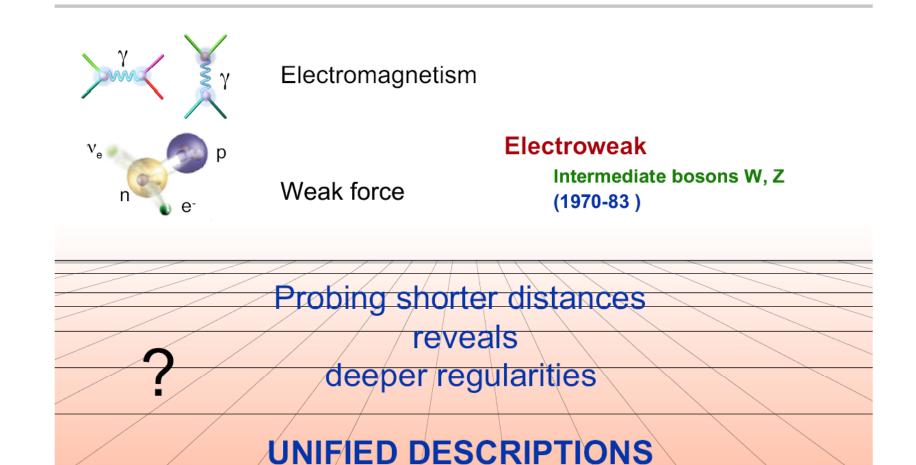
Electricity

Magnetism

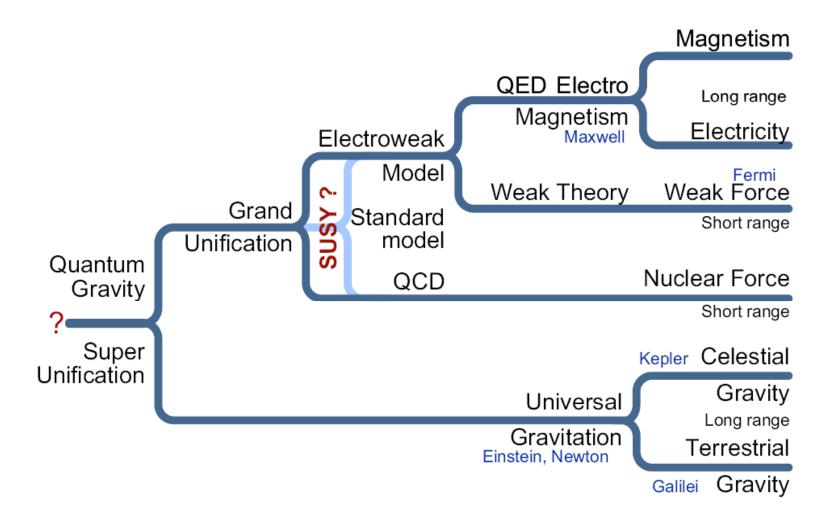
Electromagnetism

Electromagnetic waves (photon) (J.C. Maxwell, 1860)

Unification of the Forces



The Standard Model and Beyond



Quantum Gravity Era

- 10⁻⁴³ sec after big bang
- Gravity separates as a force, the other forces remain as one (Grand Unification)



- $t < 10^{-43}$ s: The Big Bang
 - Universe expands from single point with infinitely high energy density (infinite T)
- T $\approx 10^{-43}$ s, 10^{32} K
 - 10¹⁹ GeV, 10⁻²⁴ m
 - All particle types and their antiparticles in thermal equilibrium (created and annihilated at same rate)
 - Phase transition: gravity "freezes out" and became distinct from the other forces
 - Other 3 forces cannot be distinguished from one another in their interactions with quarks and leptons

Grand Unification Era

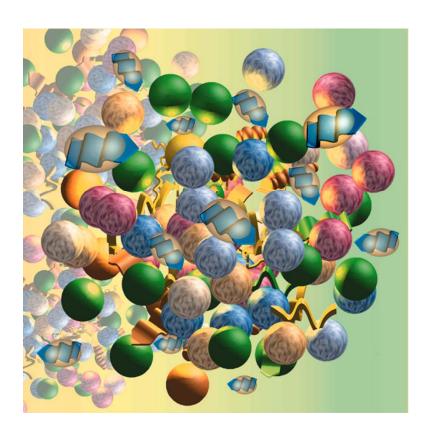
- 10⁻³⁵ sec after big bang
- Inflation ceases, expansion continues
- Strong and electroweak forces become distinguishable



- t $\approx 10^{-35}$ s, 10^{27} K: Inflation
 - 10¹⁶ GeV, 10⁻³² m
 - Rate of expansion increases exponentially for a short time
 - Universe doubled every 10⁻³² s
 - Presently observable universe only 3 m in size after inflation
 - Solves horizon and flatness problem
- T $\approx 10^{-32}$ s, Strong force freezes out
 - Via another phase transition
 - Slight excess of matter over antimatter develops (1 ppb) sufficient to yield presently observed excess of matter over antimatter
 - Too hot for quarks and gluons to bind together: they form a quark-gluon plasma

Electroweak Era

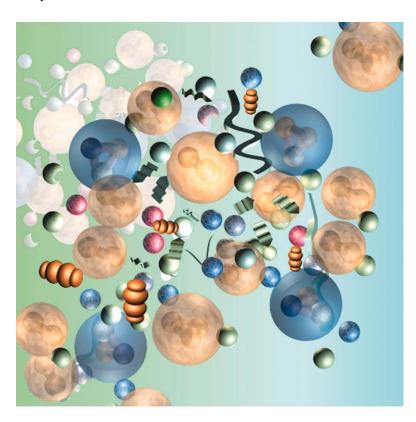
- 10⁻¹⁰ sec after big bang
- Electroweak force splits



- t $\approx 10^{-10}$ s, 10^{15} K: Electromagnetic and weak forces separate
 - 100 GeV, 10⁻¹⁸ m
 - Weak force "freezes out" via another phase transition
 - All four forces become distinct in their actions
 - Anti-quarks annihilate with quarks leaving a residual excess of matter
 - W and Z bosons decay

Protons and Neutrons Form

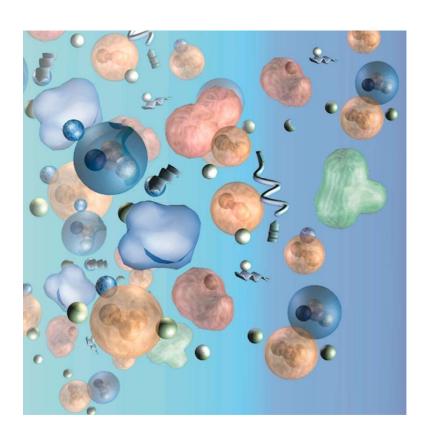
- 10⁻⁴ sec after big bang
- Quarks combine to make protons and neutrons



- t $\approx 10^{-4}$ s, 10^{13} K: Protons and neutrons form
 - 1 GeV, 10⁻¹⁶ m
 - Universe has grown to the size of our solar system
 - As T drops, quark-antiquark annihilation stops and the remaining quarks combine to make protons and neutrons
- t $\approx 1 \text{ s}$, 10^{10} K : Neutrinos decouple
 - Neutrinos become inactive (essentially do not participate in interactions)
 - Electrons and positrons annihilate and are not recreated: an excess of electrons is left

Nuclei Form

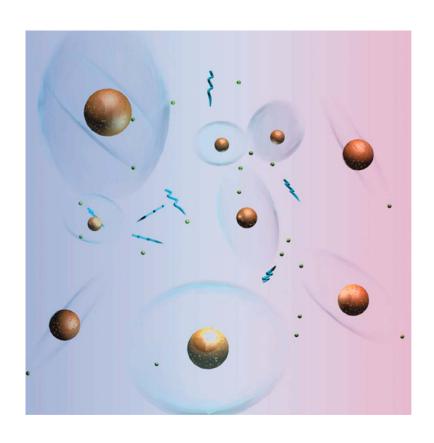
- 100 sec after big bang
- Protons and neutrons combine to form helium nuclei



- t ≈ 3 min, 10⁹ K: Nuclei are formed
 - 0.1 MeV, 10⁻¹² m
 - T now low enough to allow nuclei to form
 - Conditions are similar to those in a star today or in thermonuclear bomb
 - Neutron-proton ratio drops to 13:87
 - The bulk consitution of the universe is now in place: mostly protons (75%) and helium nuclei
 - T is still too high for atoms to form and electrons form a gas of free particles

Era of Atoms and Light

- 300,000 years after big bang
- The universe become transparent and fills with light



- t ≈ 300,000 years, 6000 K: Atoms are created
 - 0.5 eV, 10⁻¹⁰ m
 - Electrons and nuclei bind together
 - Atoms of hydrogen, helium and lithium are created
 - Radiation (light, photons) is no longer energetic enough to break atoms
 - The universe becomes transparent to photons
 - Matter density dominates
 - Astronomy/astrophysics can study the evolution of the Universe back to this time (COBE, WMAP, PLANCK)

Galaxy Formation

1000 years after big bang



- $t \approx 10^9$ years, 18 K: Galaxy formation
 - Local mass density fluctuations act as seeds for stellar and galaxy formation
 - Nucleosynthesis, the synthesis of nuclei such as carbon up to iron, starts to occur in stars
 - Heavier elements are synthesized and dispersed in stellar collapse and supernovae explosions

Today

13.7 billion years after big bang



- t $\approx 13.7 \times 10^9$ years, 3 K: Humans evolve on earth
 - Chemical process link atoms to form molecules, DNA, life
 - Humans build COBE, Hubble, Tevatron, LHC etc. to figure it all out

Credits

- CMS Brochure 2003
 - http://cmsinfo.cern.ch/outreach/CMSdocuments/CMSbrochure/CMS-Brochure03.pdf
- T. Virdee, Inaugural Lecture
 - <a href="http://cmsinfo.cern.ch/outreach/CMSdocuments/JimInaugural/J