### Dark Matter

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Teacher Academy 06/27/2014

# Physics at Different Scales

### Particle Physics



### Nuclear Physics



Condensed Matter Physics



### Cosmology

#### Astronomy



### **Biophysics**



# Large Scale Frontier

|--|--|--|--|

Stars	Galaxies	Clusters	Observed Universe
~10 <sup>11</sup> m	~10 <sup>21</sup> m	~10 <sup>23</sup> m	<b>~Ⅰ0</b> <sup>26</sup> m

# History of the Universe



• We try to understand the universe as a whole

### Evidence for Dark Matter

### • Galaxy clusters





In the 1930's Fritz Zwicky found that the galaxies in the Coma cluster were moving too fast to be contained by the visible matter

### Evidence for Dark Matter



# Evidence for Dark Matter

### • Spiral galaxies (Rotation curves of galaxies)







Vera Rubin and her collaborators (1970s)

- Expect v drops beyond luminous region
- Find v is nearly a constant
- The discrepancy is resolved by dark

#### matter

# Milky Way Dark Halo





Dark halo mass:  $\sim 10^{12} M_{sun}$ Total stellar mass:  $\sim 6 \times 10^{10} M_{sun}$ 

# Components of the Universe





local dark matter information: density: about one particle per coffee cup (if the mass is 100 times the proton mass) velocity: about 220 km/s

### Small Scale Frontier





### Size: about 10<sup>-10</sup> m

### **Fundamental Interactions**



### **Dark Matter Properties**



Not luminous Not short-lived Not hot Not baryons



The successful standard models of cosmology and particle physics are inconsistent

### Dark Matter Interactions



### Dark matter candidate:

Add a new massive particle X
Interacts with us through the weak interaction

Weakly-Interacting Massive Particle (WIMP)

# The WIMP Paradigm

### • The WIMP Miracle







### Expected event: less than I event/kg/year





### **Direct Detection** Cryogenic Dark Matter Search (CDMS)



Operating at milli-Kelvin temperatures in a mine in Minnesota

1 cm



# Annihilation Signals







# IceCube at the South Pole



Look for neutrinos from dark matter annihilation

# International Space Station



Look for electrons/photons from dark matter annihilation

### **Collider Search**



# Large Hadron Collider

L.H.C.b



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### Detector at the LHC



### More than 10,000 tonnes





# Beyond the WIMP Paradigm

- WIMPs are a good dark matter candidate
- Physicists are searching for WIMPs
- But what if dark matter is not a WIMP...



# Beyond the WIMP Paradigm

### • Hidden sector dark matter



• An example: hidden charged dark matter



# Self-interacting Dark Matter

• Self-interactions can affect dark matter dynamics



The self-interacting dark matter can behave like hot gas

randomize the dark matter velocity dispersion; lead to spherical dark halos

### Limit on the Hidden Charge



# Summary

- We have two successful theories of the large and small, but they are not consistent
- A quarter of the universe is dark matter, but we don't know what it is
- We have some ideas and many search experiments are underway